



ABORT program on Fatal Error.



Kavli Institute
for Cosmological Physics
AT THE UNIVERSITY OF CHICAGO

SNANA Tutorial

R. Kessler

Duke Seminar
June 18, 2020



From the SNANA Legal Team:

- **SIDE EFFECTS** include, but are not limited to: confusion, frustration, watery eyes, headaches, weight loss, weight gain, systematics-limited results, incorrect results, denial of tenure.
- Do not drive or operate heavy machinery while using SNANA.

Snana

Y A T R A



GOOGLE Search: No, not this SNANA



Welcome to the SuperNova ANAlysis software homepage

- Install Guide
- SNANA Manual
- Overview Paper
- Legal Notice
- SNANA Download**
- Mailing List
- Contact Form

SNANA contains a light curve fitter and simulation that can be applied to any supernova (SN) model and to any data set. This website provides installation instructions, a user manual, and a software package download area.



And not this
obsolete link

It's here on Github:
<https://github.com/RickKessler/SNANA>

RickKessler / SNANA

Watch 4 Star 8 Fork 8

<> Code Issues 2 Pull requests 1 Actions Projects 0 Wiki Security 0 Insights Settings

Supernova Analysis package Edit

Manage topics

659 commits 16 branches 0 packages 43 releases 6 contributors

Branch: master New pull request Create new file Upload files Find file Clone or download

RickKessler misc	Latest commit 8631d18 34 minutes ago
doc	misc 34 minutes ago
src	implement +64 bit for SPECTROGRAPH_OPTMASK 34 minutes ago
util	fix abort message for missing datafile= key 34 minutes ago
.gitignore	updated gitignore 11 months ago
README.md	Merge remote-tracking branch 'origin/master' into BYOSED 13 months ago

README.md

SNANA

Supernova Analysis package.

Read documentation in /doc.

Outline

- Intro
- Architecture
- Main programs and features
- Systematics & Multi-Core Jobs
- Pippin Analysis Pipeline (Super Glue)
- Output
- Documentation

Why SNANA

- ***SN Ia-Cosmology Analysis***
(SDSS, JLA, PS1-Pantheon, DES)
- Forecast future surveys (DES, LSST, WFIRST ...)
- Simulate wide range of transients
(e.g., PLAsTiCC classification challenge)
- Quantify sensitivity for rare searches
(e.g., KN, ICECUBE- ν source)
- Efficiency correction for rate measurements
- Validate diff-imaging pipeline (using fakes)

Few Principles for SNANA

- ***Science needs drive development***
- Stay connected to real data, even if designing a future survey
- Always look for opportunities to try new things, even if it seems like a distraction.
- Minimize memory, process time, file size.

Summary of Ready-to-Run Programs in \$SNANA_DIR

- Simulation package (catalog, not pixels)
- Light Curve Fitting & Template Fitting
- Bias Corrected Hubble diagram using BBC
- Create stat+syst covariance matrix
- Cosmology Fitting (old, simple, fast: better codes elsewhere)
- Utilities for systematics & multi-core processing
- *NO Image-Processing Tools*

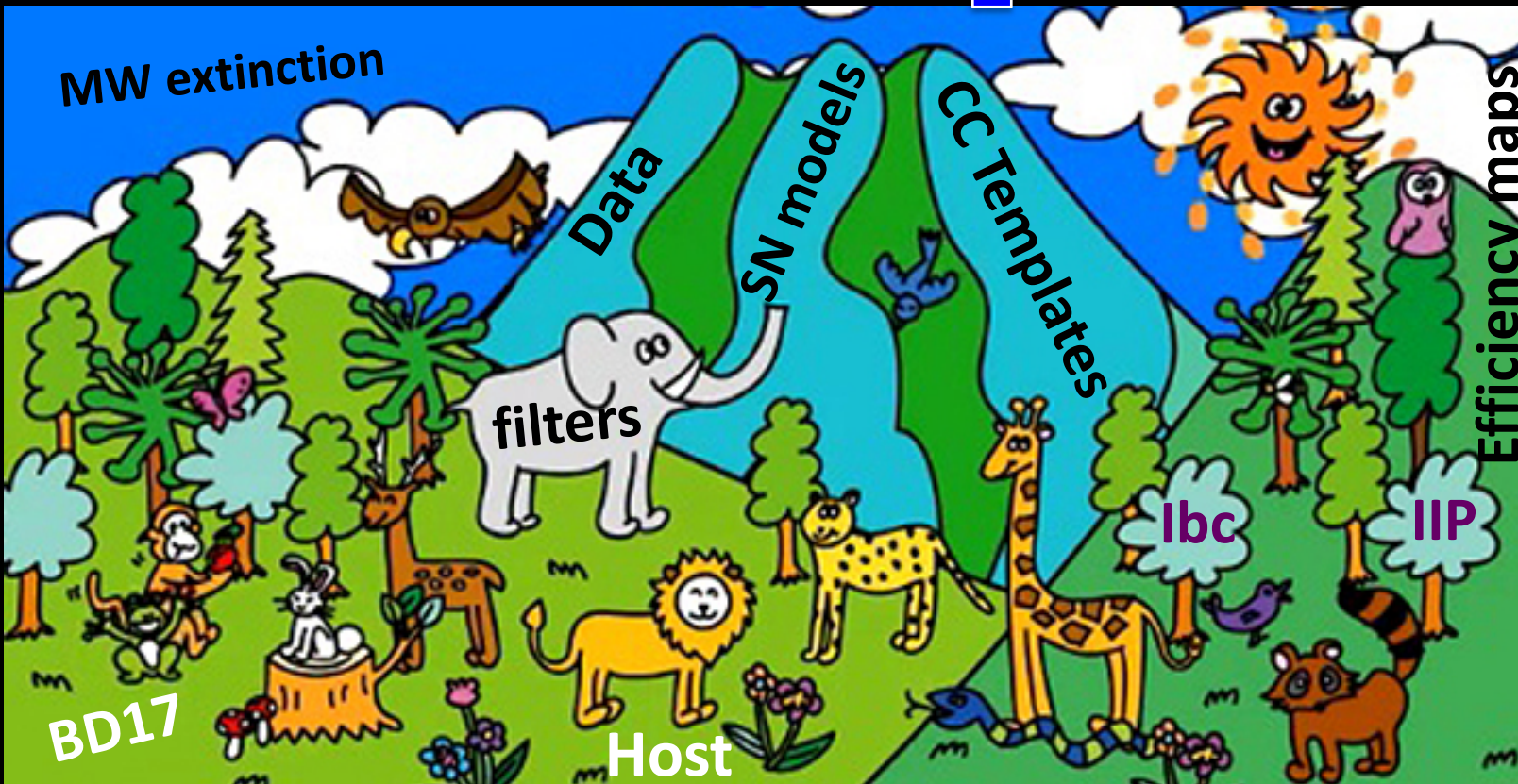
Code Languages

- sim, BBC, wFit: C
- LC fit: C + Fortran (unintended consequence)
- Batch-submit scripts: Perl (naive mistake)
- createCov + plot-util + misc: Python

Architecture: Environment

- Simulation package (catalog, not pixels)
- Light Curve Fitting & Template Fitting
- Hubble Diagram
- Utilities for systematics & multi-core processing

`$SNANA_DIR`



`$SNDATA_ROOT`

SNANA Architecture: File Sharing

SNANA was implicitly designed to run on a cluster with many users sharing files.



Hey, did you update the filter transmission files ?

SNANA Architecture: File Sharing

SNANA was implicitly designed to run on a cluster with many users sharing files.

`$SNDATA_ROOT` contains

- Data
- Simulated output
- Filter transmissions
- Primary SEDs
- Ia & CC spectral templates
- MW extinction map
- SN model parameters
- SIMLIB files
- HOSTLIB files
- Efficiency maps
- Etc . . .

Can run tests with file(s) in your private directory, but goal is to share files with *community* via `$SNDATA_ROOT`

SNANA Architecture: Sharing Proprietary Files

During analysis it is useful to share proprietary SNANA files,

\$DES_ROOT

\$LSST_ROOT

\$WFIRST_ROOT

INTERNAL directories are *NOT* public.

ENV can be used as part of any input **fileName**, e.g.,

SIMLIB_FILE: \$DES_ROOT/simlibs/DES_DIFFIMG.SIMLIB

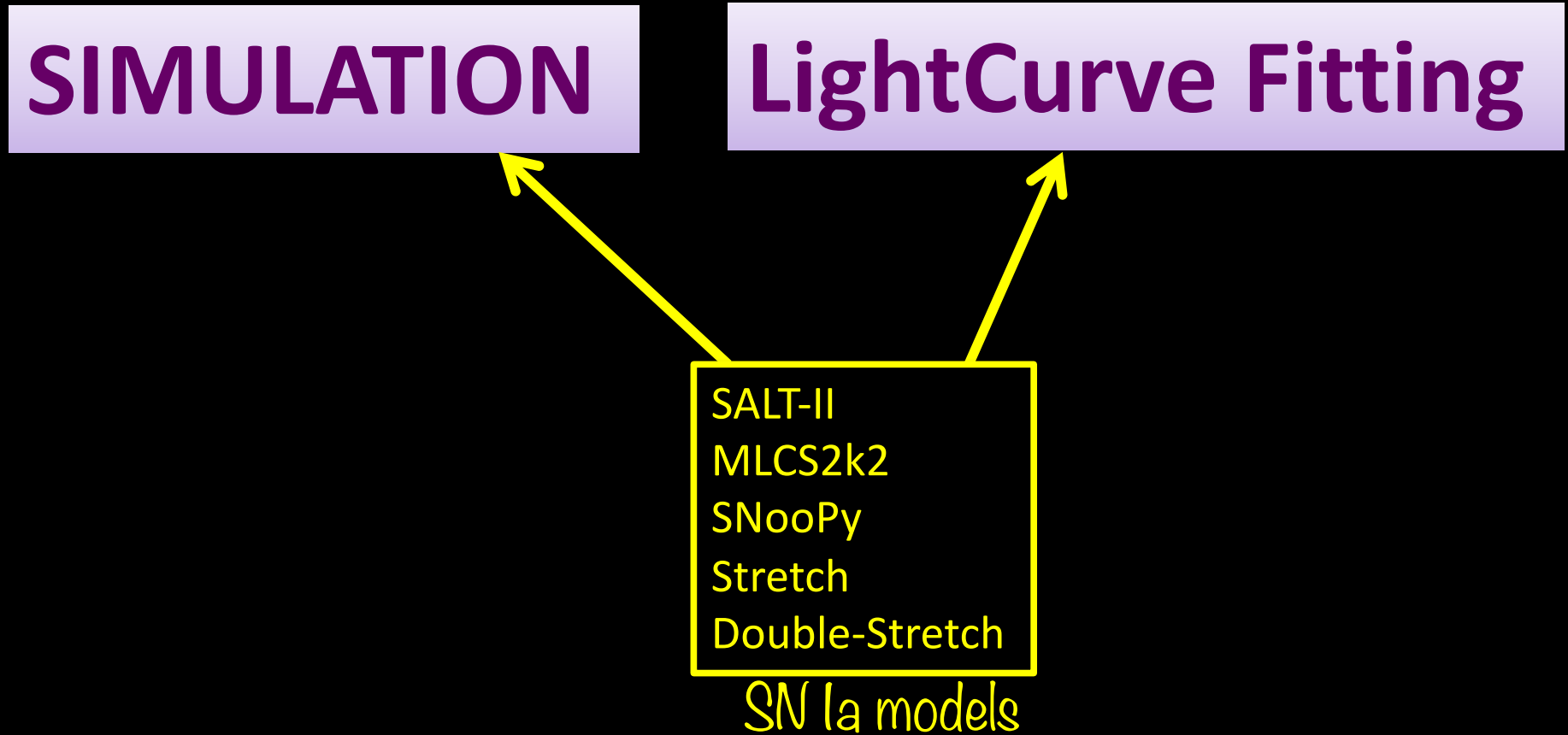
(ENVs simplify transfer to another compute cluster)

Surveys Investigated on U.Chicago's "Midway" Cluster

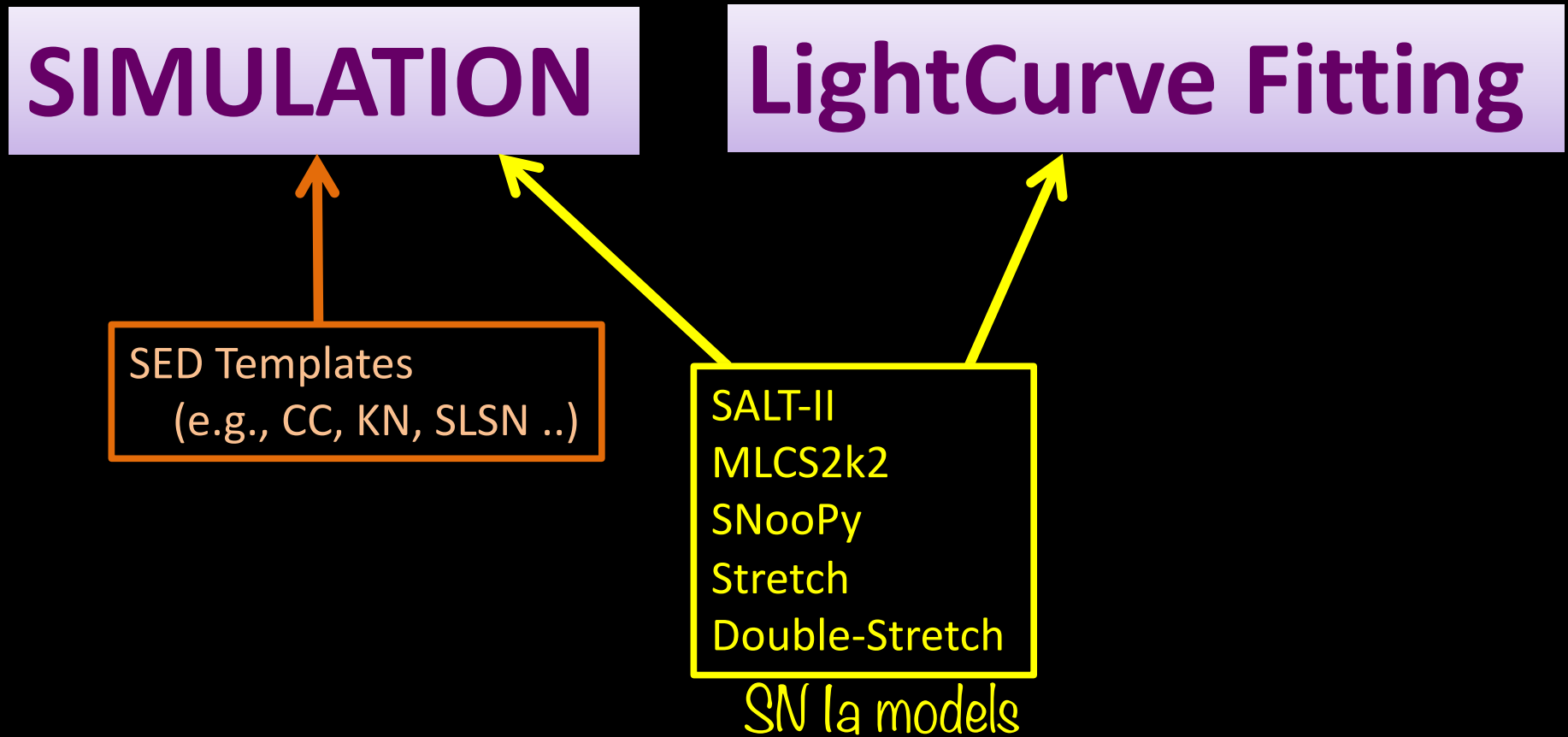
DES JLA PS1MD SMT WFIRST ZTF
FOUNDATION LSST SDSS SNTRAIN YSE

- No barriers between project areas
- Easy to switch between projects (no new login)
- Easy to collaborate with others
- Relies honor system to respect proprietary files

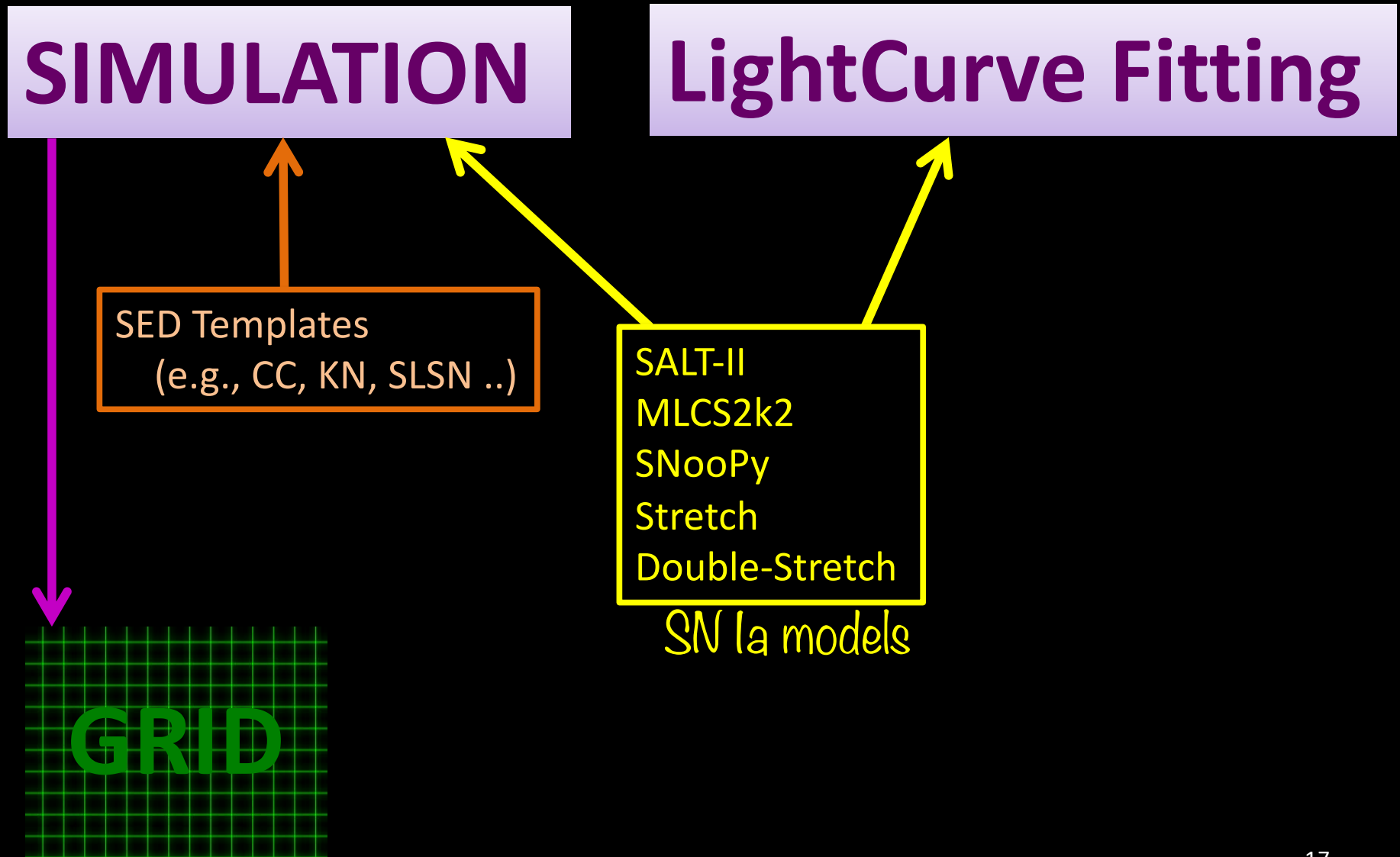
SNANA Architecture



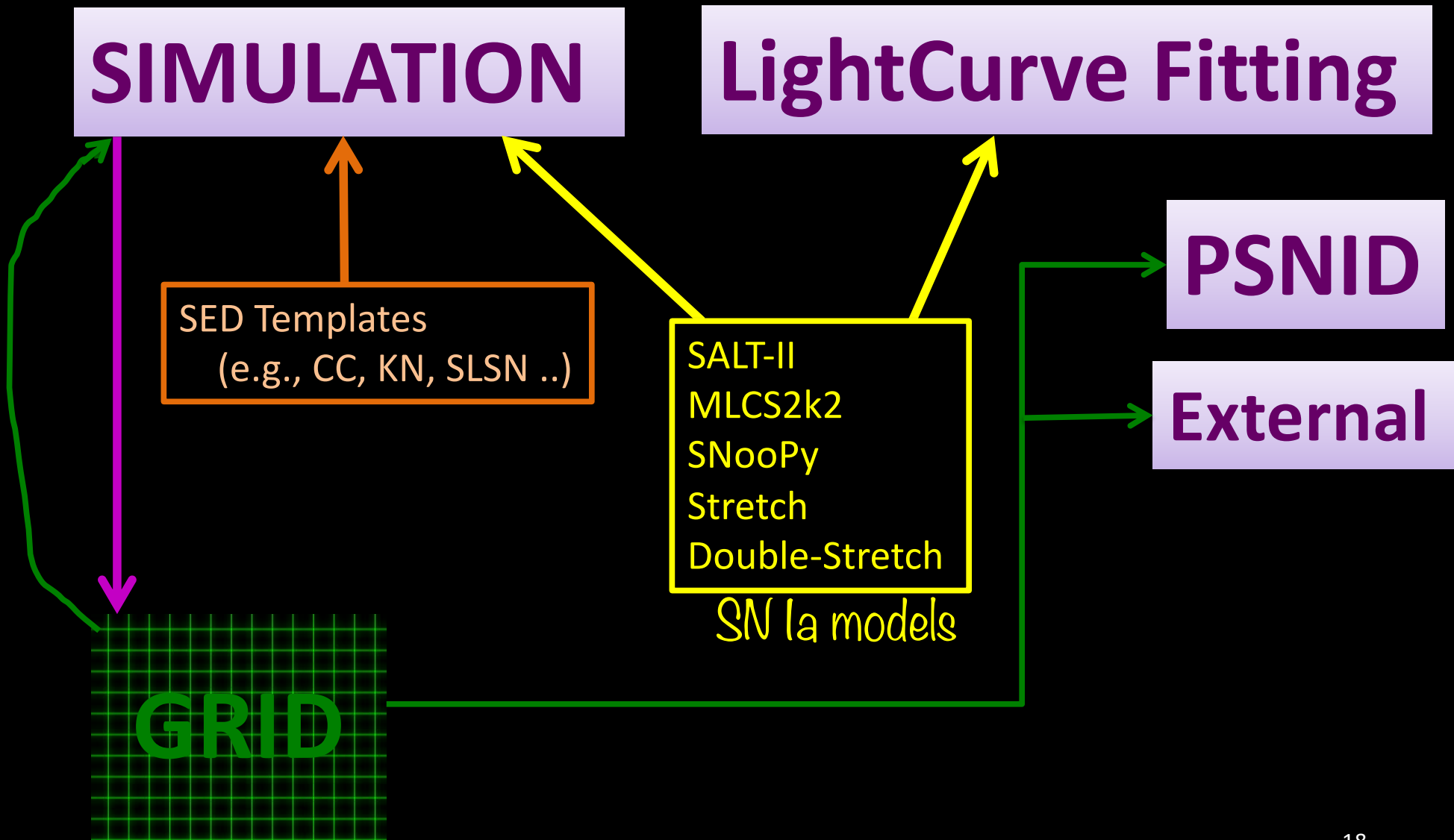
SNANA Architecture



SNANA Architecture



SNANA Architecture



Why Use GRID ?

(versus model params & redshift)

- Allows using templates constructed from non-SNANA programs
- Any SN model \rightarrow standard model format for template-fitting programs such as PSNID
- Can be faster (e.g., huge speed-up for SNooPy)

SNANA Simulation

SOURCE MODEL

Explosion Model
(e.g., Woosley, Kasen, FLASH)

Empirical model from Data (e.g., SALT2, SED time series ...)

SIM

INSTRUMENTAL MODEL from IMAGES

- Cadence
- PSF
- ZP
- Skynoise
- Non-Poisson noise
- Saturation
- Detect efficiency

**Calibrated light curves
and uncertainties**

Source Model



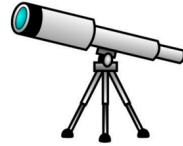
Generate source SED
(each epoch)

Cosmological dimming,
lensing, pec. velocity,
redshift SED

Galactic extinction

Integrate redshifted
SED for each filter

Noise Model



Use cadence info
(PSF, SKY, Zeropoint)
to convert true mag
to true flux (ADU)
and true uncertainty

Apply Poisson noise
to get measured flux.

Trigger Model



Check for detection

Apply candidate logic:
e.g., 2 detections

Apply spectroscopic
selection function

Write selected events
to data files.

Almost 20 different “maps”
are input to the simulation

Host Galaxy Library (HOSTLIB)



- mis-matched host based on min DLR
- SN-host correlations (e.g., brightness vs. mass)
- SN population (color, stretch, dust) dependence on HOST properties (e.g, mass, SFR ...)
- Host photo-z (e.g., from Graham 2018)
 - use as prior in fit for SN photo-z
- Local surface brightness \rightarrow Poisson & non-Poisson (anomalous) noise.
- Efficiency map of spec-zHOST vs. HOST properties

Host Galaxy Library (SN-Host correlations)



- Given SN properties (color, stretch),
choose distribution of host properties
or
- Given host properties, choose
distributions of SN properties.
(color, stretch, RV, AV)

Host Galaxy Library

(Example HOSTLIB using MICECAT)



```
VARNAME: GALID RA_GAL DEC_GAL ZTRUE  
g_obs r_obs i_obs z_obs  
ABS_MAG_g ABS_MAG_r ABS_MAG_i ABS_MAG_z  
a0_Sersic b0_Sersic n0_Sersic a_rot  
LOGMASS LOG_SFR
```

```
GAL: 28581888 6.32295 25.82068 0.30917  
22.47973 21.19873 20.77373 20.56673  
-18.71227 -19.31427 -19.55727 -19.76027  
0.91923 0.36473 2.0 -162.357 9.89 -10.2
```

etc ...

Host Galaxy Library (HOSTLIB)



WARNING

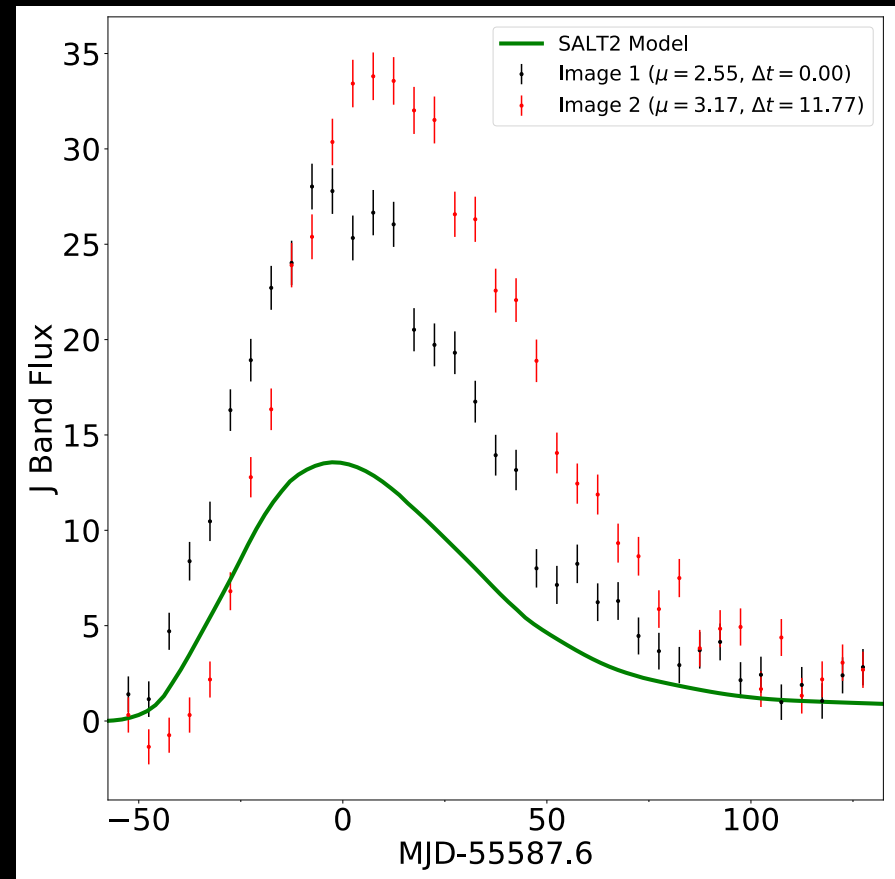
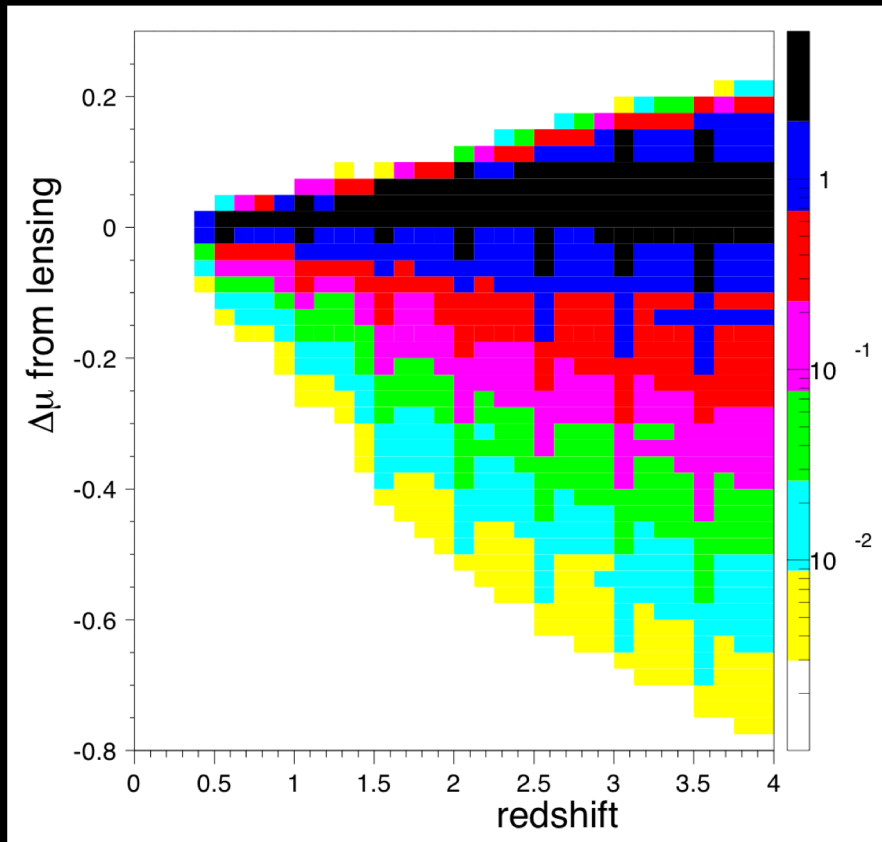
SNANA code implementation is the easy part ... main burden is on the user to construct suitable HOSTLIB

Simulation of Weak & Strong Lensing

Nancy Grace Roman
Space Telescope

J. Pierel et al, in prep

Weak lens magnification map from
J.Barreiro, T. Davis (MICECATv1)



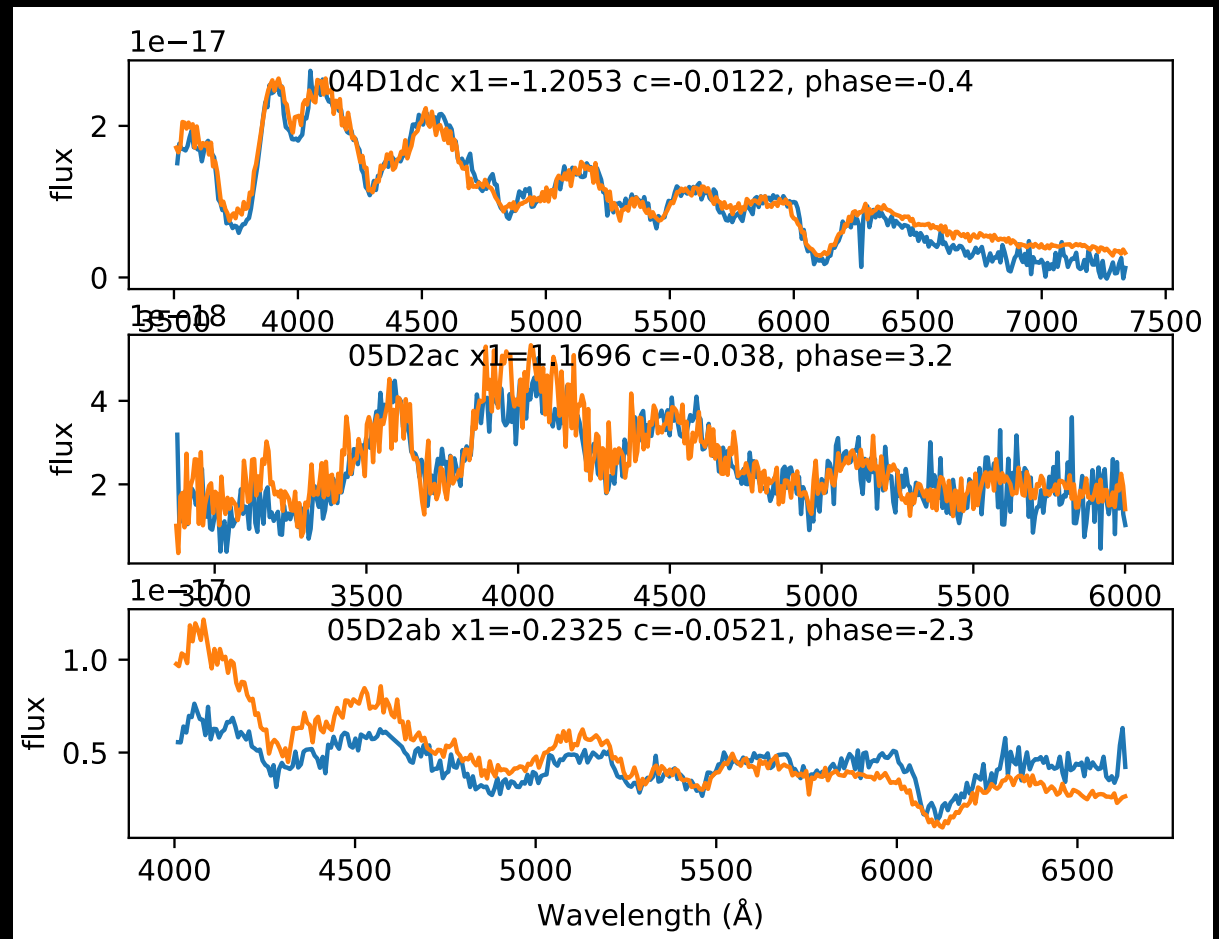
Simulating Spectra: SN & Host

- Test SALT2 training with simulated photometry & spectra
 - Impact on spectral calibration
 - Impact of host contamination
 - Impact of spectral wave & phase coverage
- NGRST (WFIRST): forecast host redshift capabilities

Simulating Spectra: SN

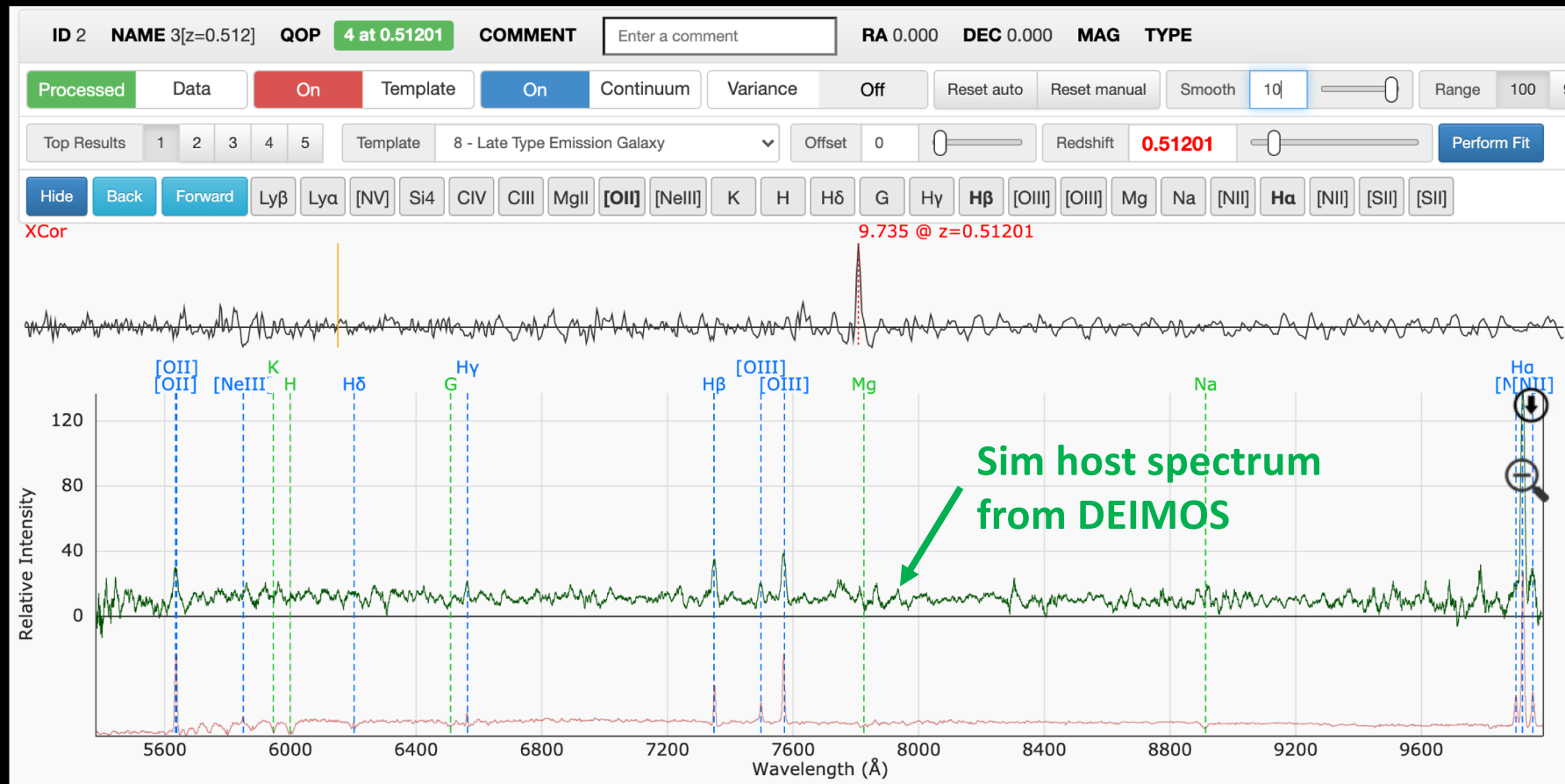
--- Real SNIa, SNLS

--- Sim SNIa, SNANA



Plot from D.Jones

Simulating Spectra: Host (and reformat option for Marz redshift utility)



- Spectrograph Sim from S.Rodney
- <http://samreay.github.io/Marz> (by S.Hinton)

SIM CPU Proc-Time

U Chicago Research Computing Center:
Sep 2018 for PLAsTiCC

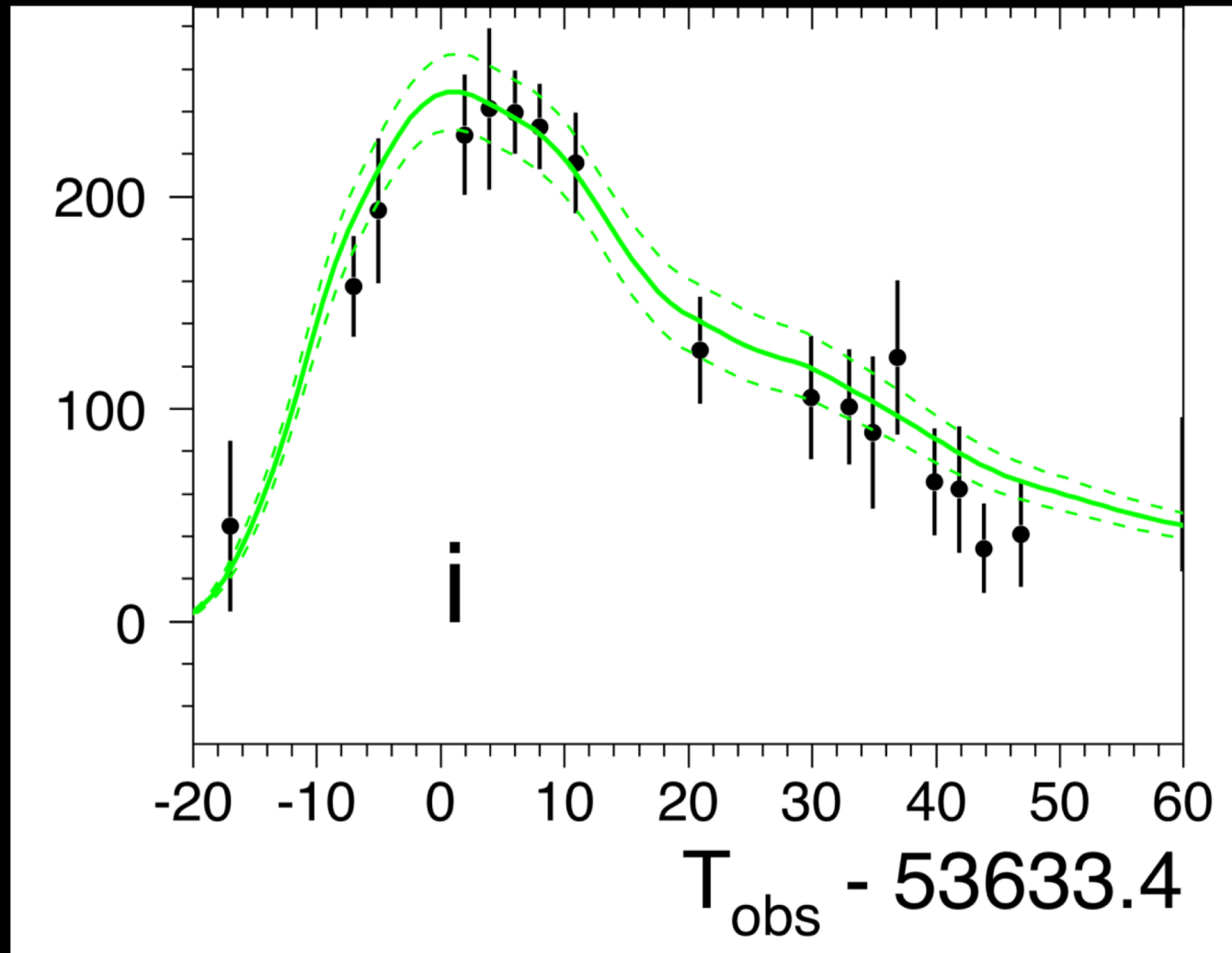


- 117 million light curves generated in 8 hr on 40 cores → 100/sec
- 3.5 million written to data file → 3/sec

No API for Simulation

- However, without re-compiling can add
 - + new survey, filters, Cadence, HOSTLIB, calibration info
 - + new SED templates
- Supernova Standardization Team (SST) has added ``Build Your Own SED'' model (BYOSED) in python, with C calling python.
- SST Goal: arbitrary complexity in sim data to test SALT2 training.

Analysis and Light Curve Fitting

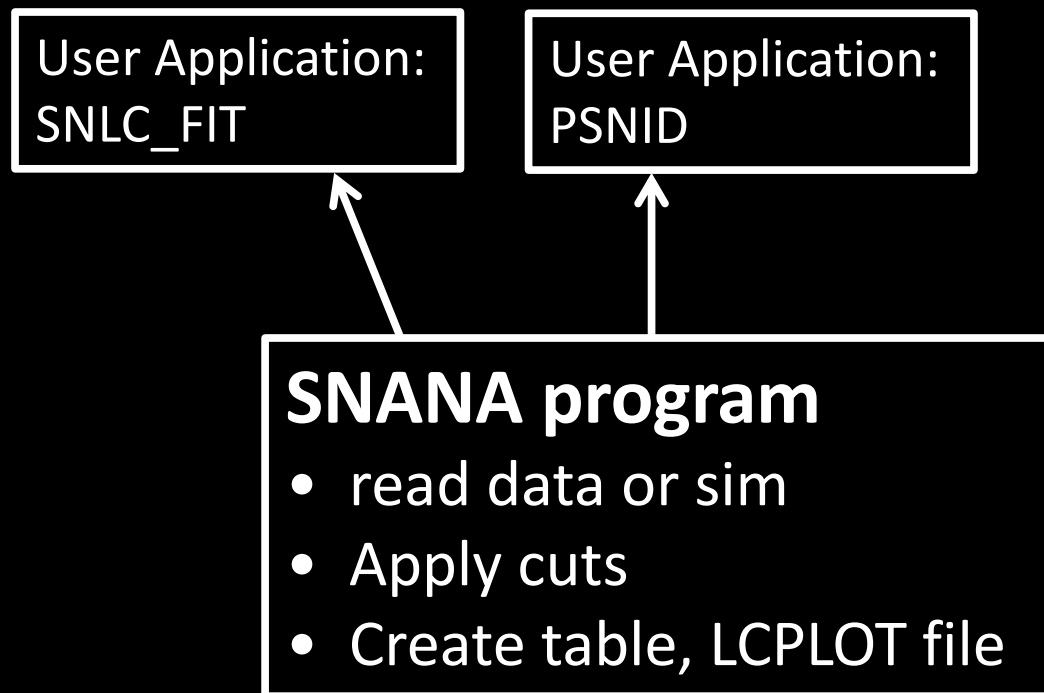


Architecture for Analysis

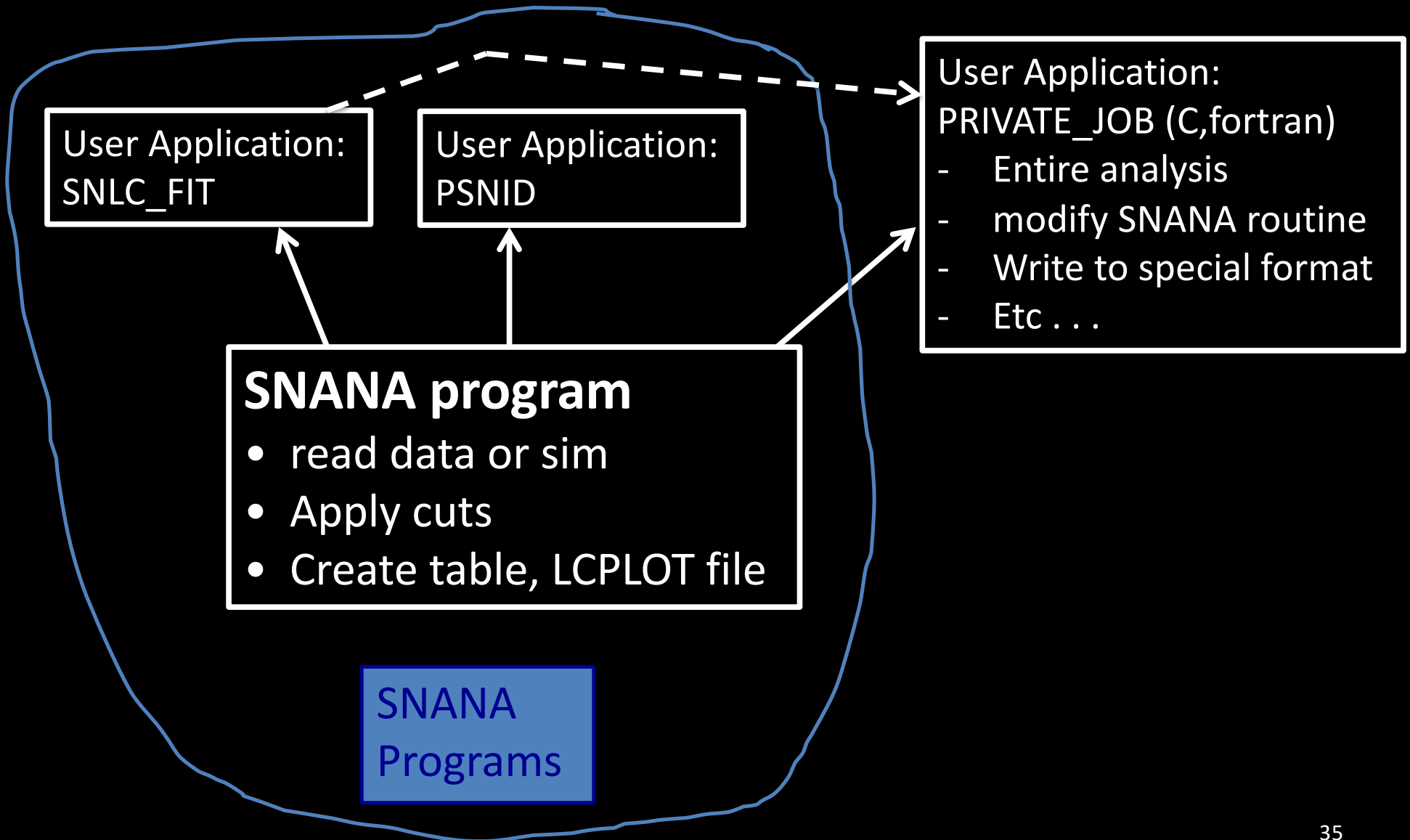
SNANA program

- read data or sim
- Apply cuts
- Create table, LCPLLOT file

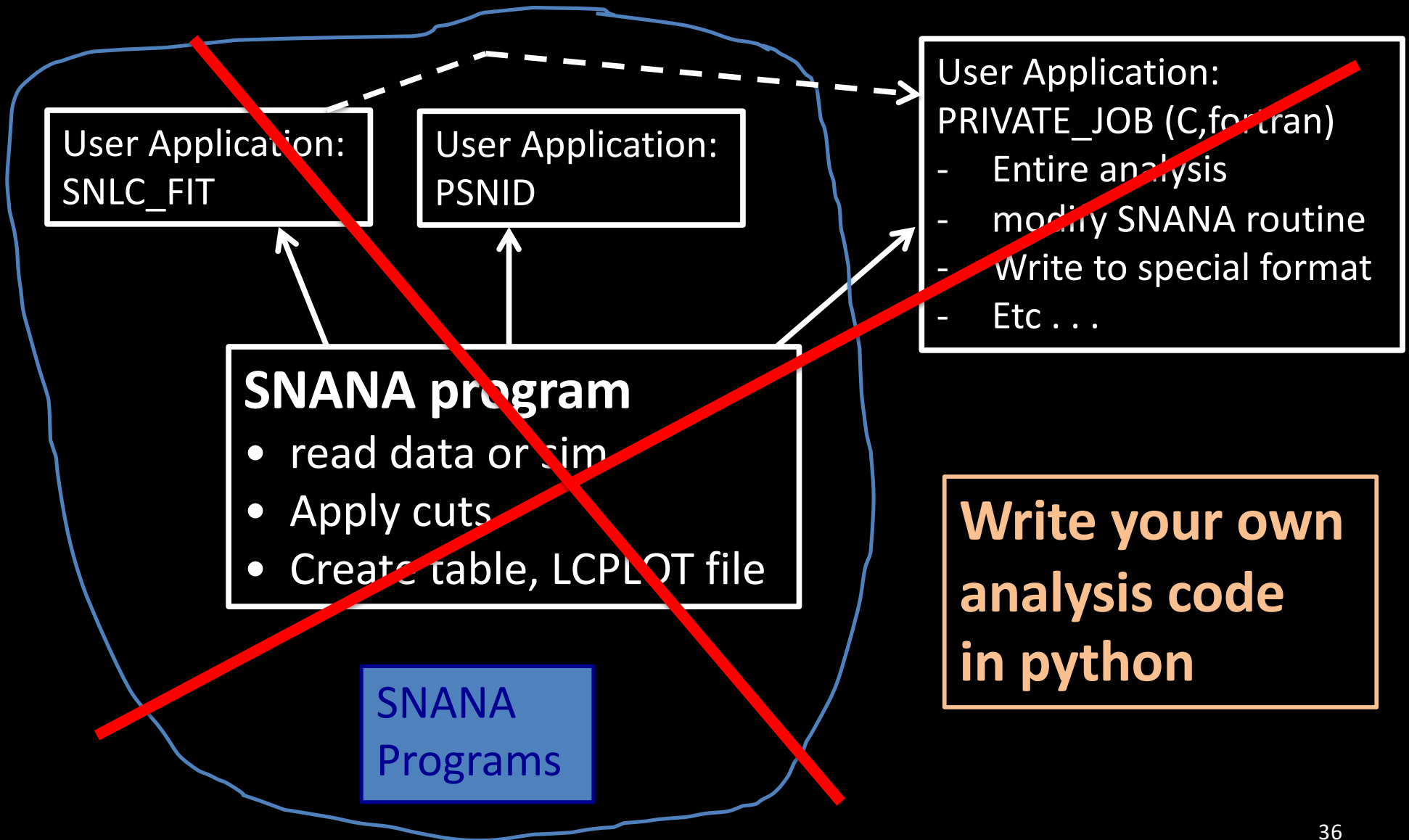
Interface for Analysis



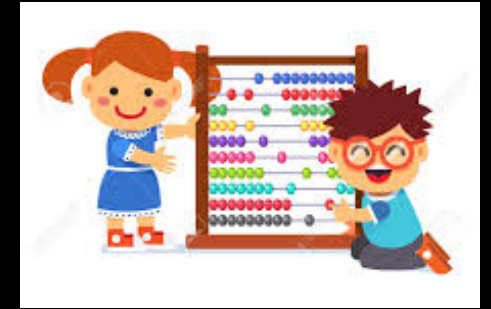
Interface for Analysis



Interface for Analysis



Misc Analysis Options



- ``CUTWIN'' options to apply selection requirements
- Systematic variations (e.g., calibration, MW-extinct, bands..)
- Override data variables with supplement file (e.g., try out different pec. velocities)
- Apply fluxErr map to scale uncertainties

BEAMS with Bias Correction (BBC)

SALT2 fit params from:

- Real data
- Simulated data (Ia)
- Simulated data (CC)

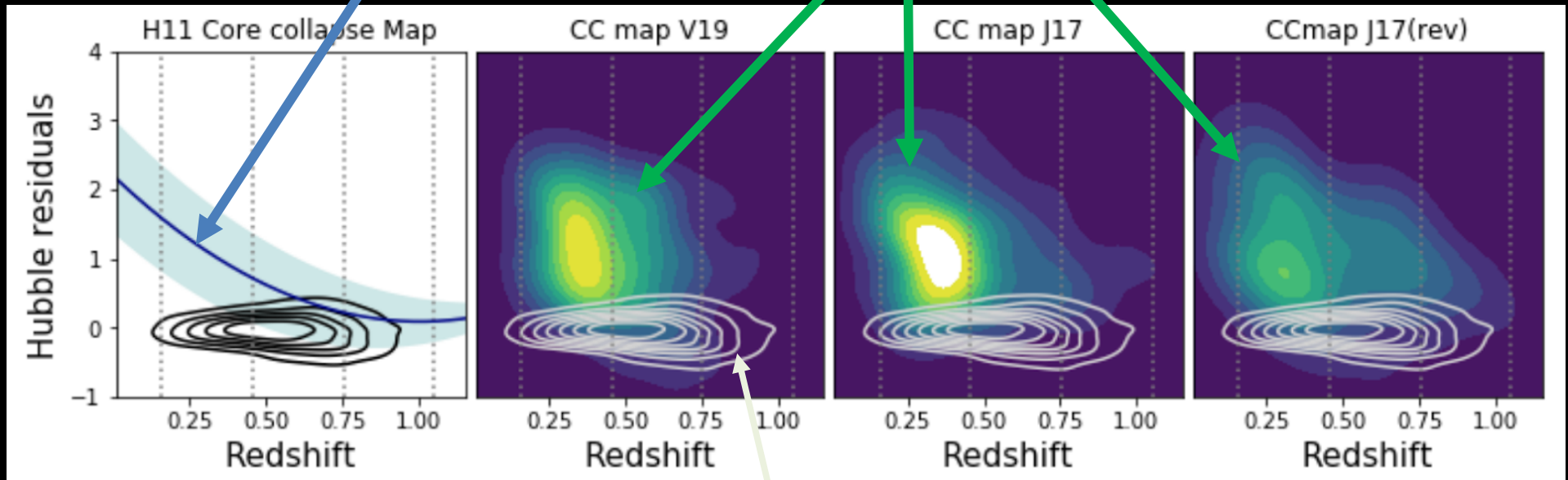
BBC fit
program

z-binned Hubble diagram
corrected for bias from
selection and CC contam.

BEAMS with Bias Correction (BBC)

Analytic CC prior
(Hlozek 2011)

Simulated
CC prior:



Plot courtesy of M.Vincenzi

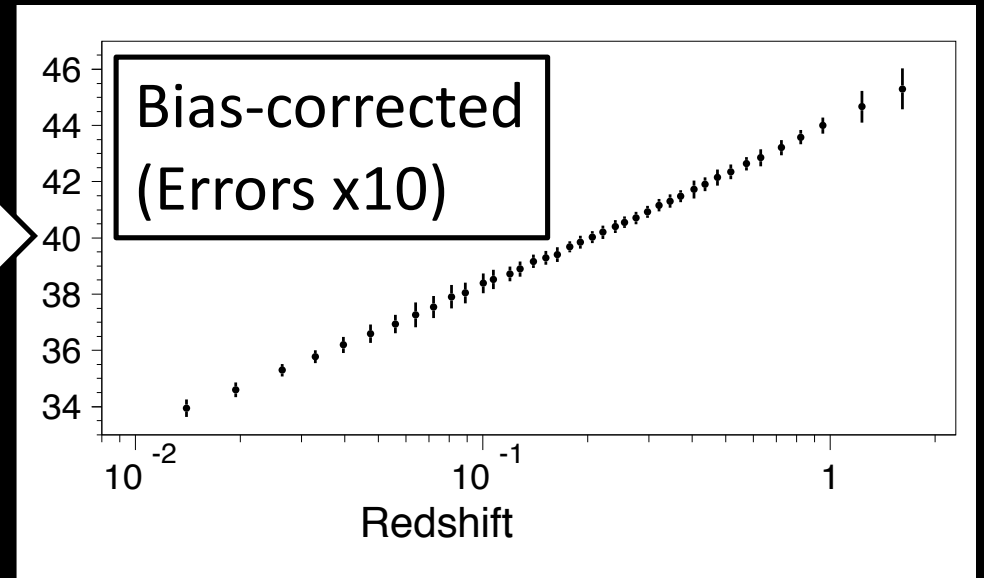
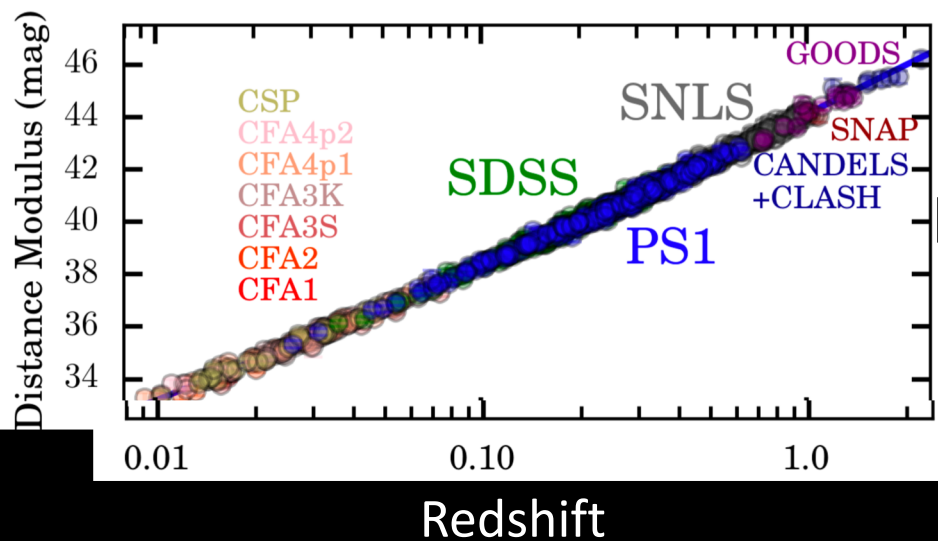
SNIa prior
(analytic)

BBC Output

(input to Cosmology fit)



Pantheon Sample (Scolnic 2018)

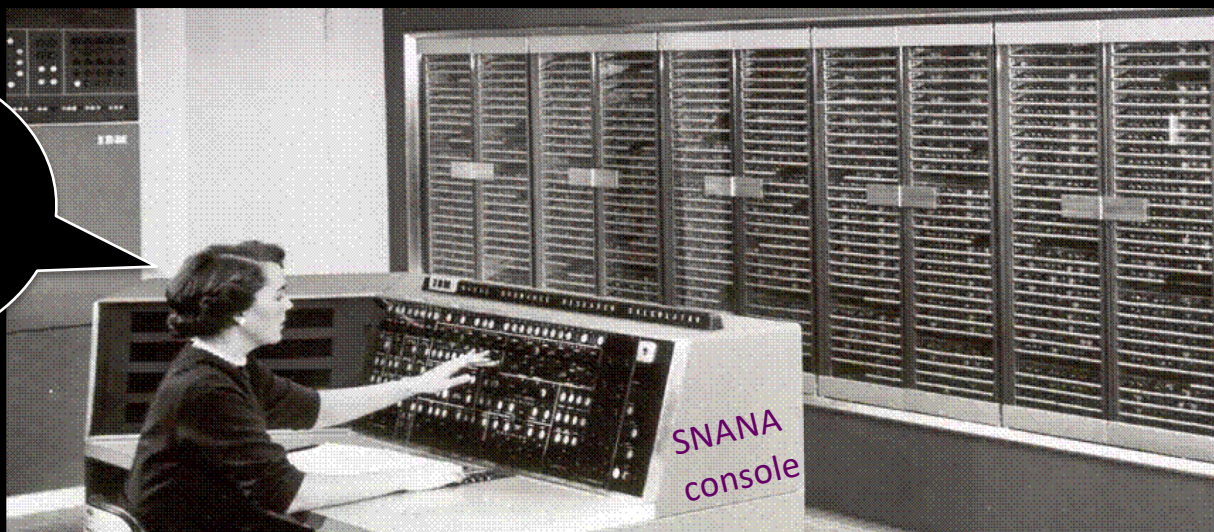


Systematics & Multi-Core Jobs

- Precision analyses require many iterations of analysis, each with a small variation in parameters or method.
- SNANA has tools to implement multi-iteration analyses using multi-core platforms.

(e.g., Fermilab, NERSC, Argonne, Midway, Folio . . .)

Switching
to C11
Intrinsic
scatter
model



Systematics & Multi-Core Jobs

- Sim & Analysis codes read input instructions from a text file.
- Specify variations with **command-line overrides** to avoid more input files.
- SNANA Scripts use **command-line override** feature to launch multiple jobs in batch system.

sim_SNmix.pl

Launch multiple
SIM(la+CC) jobs

split_and_fit.pl

Launch multiple
lightcurve fit jobs

SALT2mu_fit.pl

Launch multiple
BBC fit jobs

Systematics & Multi-Core Jobs

```
# vary intrinsic scatter models
GENVERSION: JLA_SDSS3year_G10smear
GENOPT: GENMAG_SMEAR_MODELNAME G10
GENOPT: SEARCHEFF_SPEC_FILE SPECEFF_SDSS/SEARCHEFF_SPEC_SDSS_G10smear.DAT

GENVERSION: JLA_SDSS3year_COHsmear
GENOPT: GENMAG_SMEAR 0.13 GENMAG_SMEAR_MODELNAME NONE
GENOPT: SEARCHEFF_SPEC_FILE SPECEFF_SDSS/SEARCHEFF_SPEC_SDSS_COHsmear.DAT

GENVERSION: JLA_SDSS3year_C11-0smear
GENOPT: GENMAG_SMEAR_MODELNAME C11_0
GENOPT: SEARCHEFF_SPEC_FILE SPECEFF_SDSS/SEARCHEFF_SPEC_SDSS_C11-0smear.DAT

GENVERSION: JLA_SDSS3year_C11-1smear
GENOPT: GENMAG_SMEAR_MODELNAME C11_1
GENOPT: SEARCHEFF_SPEC_FILE SPECEFF_SDSS/SEARCHEFF_SPEC_SDSS_C11-1smear.DAT

GENVERSION: JLA_SDSS3year_C11-2smear
GENOPT: GENMAG_SMEAR_MODELNAME C11_2
GENOPT: SEARCHEFF_SPEC_FILE SPECEFF_SDSS/SEARCHEFF_SPEC_SDSS_C11-2smear.DAT

GENVERSION: JLA_SDSS3year_NOsmear
GENOPT: GENMAG_SMEAR_MODELNAME NONE
GENOPT: SEARCHEFF_SPEC_FILE SPECEFF_SDSS/SEARCHEFF_SPEC_SDSS_NOsmear.DAT
```

Example:
Subset of sim
jobs for JLA
systematics.
Each job →
separate core

sim_SNmix.pl
Launch multiple
SIM(la+CC) jobs

Systematics & Multi-Core Jobs

```
FITOPT: MAGOBS_SHIFT_ZP 'g .01'  
FITOPT: MAGOBS_SHIFT_ZP 'r .01'  
FITOPT: MAGOBS_SHIFT_ZP 'i .01'  
FITOPT: MAGOBS_SHIFT_ZP 'z .01'  
FITOPT: FITMODEL_NAME 'SALT2.JLA_systematic/sys0'  
FITOPT: FITMODEL_NAME 'SALT2.JLA_systematic/sys1'  
FITOPT: FITMODEL_NAME 'SALT2.JLA_systematic/sys2'  
FITOPT: FITMODEL_NAME 'SALT2.JLA_systematic/sys3'  
FITOPT: FITMODEL_NAME 'SALT2.JLA_systematic/sys4'  
FITOPT: MWEBV_SCALE 1.10 MWEBV_SHIFT 0.00  
FITOPT: MAGOBS_SHIFT_ZP 'g 0.0024 r -0.003 i -0.008 z -0.013'  
FITOPT: MAGOBS_SHIFT_PRIMARY 'g 0.00 r -0.00 i -0.00 z 0.0'  
FITOPT: MAGOBS_SHIFT_PRIMARY 'g -0.0008 r -0.011 i -0.0051 z 0.016'  
FITOPT: MAGOBS_SHIFT_PRIMARY 'g -0.011 r -0.0053 i 0.0014 z 0.0056'  
FITOPT: MAGOBS_SHIFT_PRIMARY 'g -0.0037 r -0.0066 i -0.0043 z 0.008'
```

Example:
Subset of fit
jobs for PS1
systematics

split_and_fit.pl

Launch multiple
lightcurve fit jobs

Systematics & Multi-Core Jobs

```
# define Light curve fit output from split_and_fit
INPDIR+: $DES_ROOT/analysis/DES_LCFITS
INPDIR+: $DES_ROOT/analysis/LOWZ_LCFITS
INPDIR+: $DES_ROOT/analysis/FOUND_LCFITS

# define BBC variations
MUOPT: simfile_biascor=DES_variant1.FITRES,LOWZ_variant1.FITRES
MUOPT: simfile_biascor=DES_variant2.FITRES,LOWZ_variant2.FITRES
MUOPT: nbin_logmass=2  nzbin=16
MUOPT: CUTWIN x1 -2 2
etc ...
```

SALT2mu_fit.pl

Launch multiple
BBC fit jobs

Systematics Summary

- Systematics are expressed via
 - GENOPTs (sim biasCor variations)
 - FITOPTs (LC fit variations)
 - MUOPTs (BBC variations)
- Creating and testing these ~100 variations is a major part of SNIa-cosmology analysis.

Stat + Syst Covariance Matrix (createCov.py)

$$\tilde{X}_{NV} = \begin{bmatrix} X_{11} & \dots & X_{1V} \\ \vdots & \ddots & \vdots \\ X_{N1} & \dots & X_{NV} \end{bmatrix}$$

- After LC fit and BBC, create `cov(stat+syst)` for CosmoMC (cosmology fit program)
- Includes systematics from fit options (FITOPT) and BBC options (MUOPT)
- Allows arbitrary error-scaling for each systematic, or each class of systematic.
- Internal bookkeeping for ~ 100 systematics.

SNANA Output

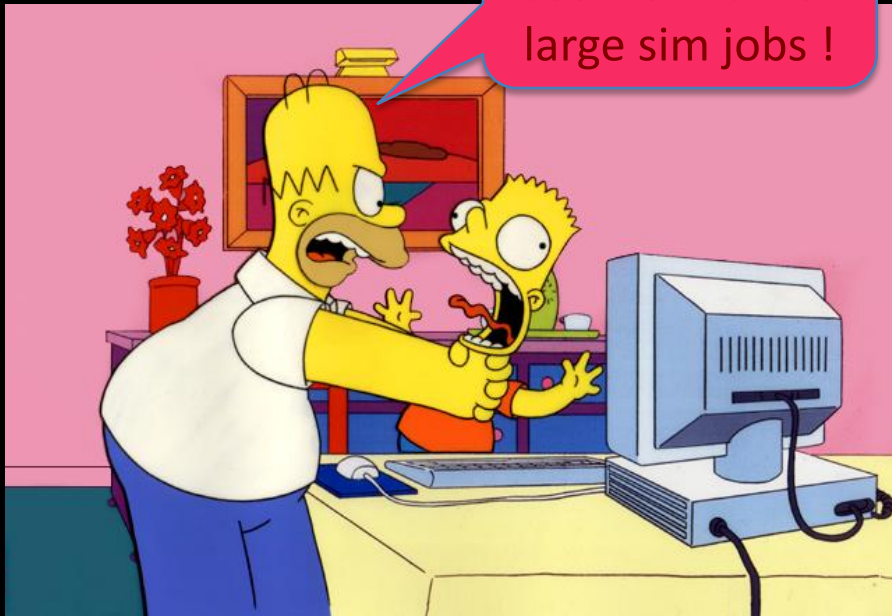


SIMULATION Output

Data Files

- FITS format for large jobs
or
- ASCII format for testing
(1 file per SN)

I said NEVER use
ascii format for
large sim jobs !



SIMULATION Output

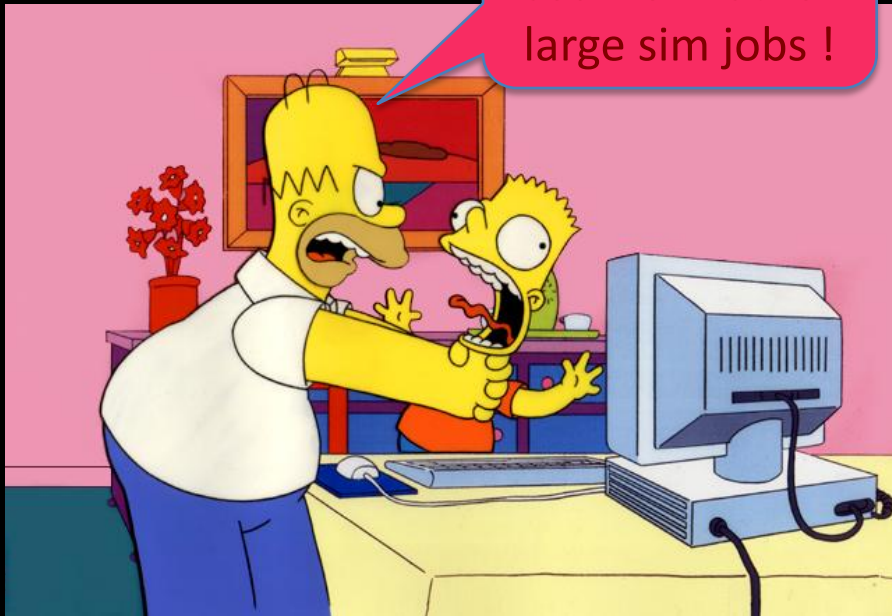
Data Files

- FITS format for large jobs
or
- ASCII format for testing
(1 file per SN)

ASCII Summary File

- 1 row per accepted SN
or
- 1 row per generated SN
(for efficiency)

I said NEVER use
ascii format for
large sim jobs !



SIMULATION Output

Never Trust Simulation Output

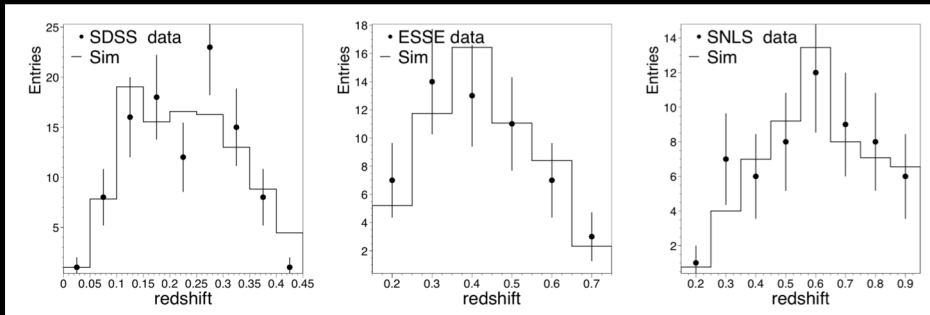


Always check data/MC distributions
(redshift, fit params, SNR, etc . . .)

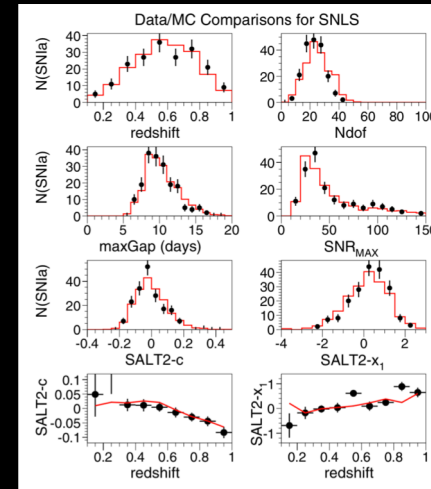


Y'all don't have to worry
none; I generated them
simulations myself.
Who took my beer ?

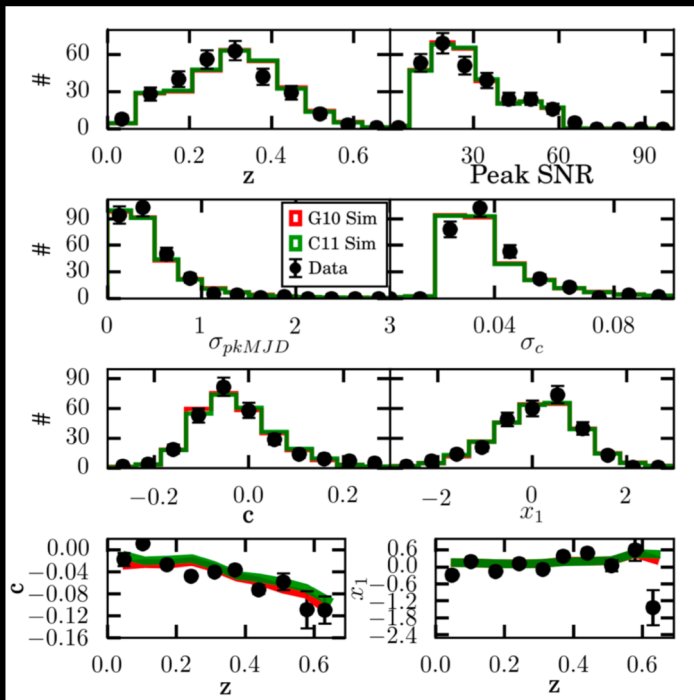
Simulation Output vs. Data



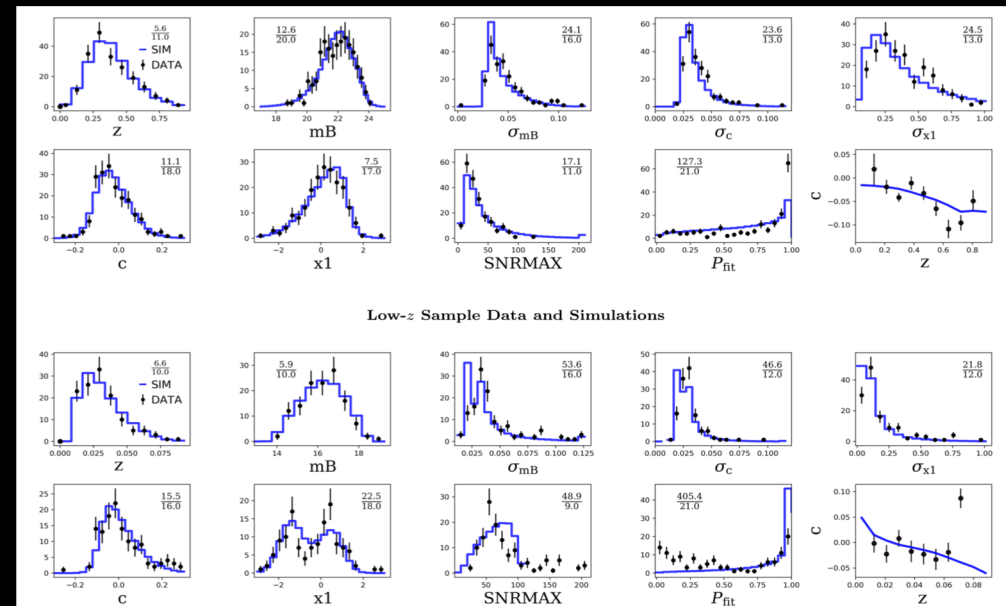
SDSS (Kessler et al 2009)



JLA
2014

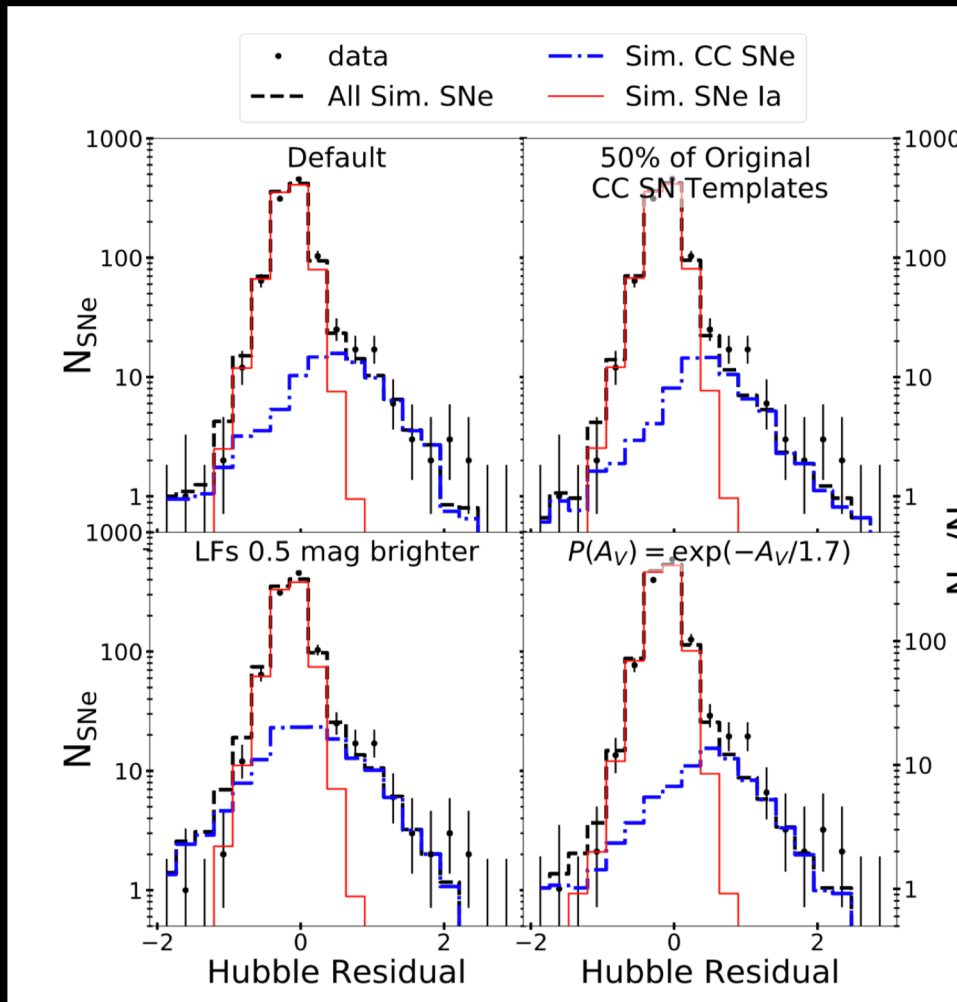


Pantheon (Scolnic et al 2018)



DES 3-Year (Brout et al 2019)

Simulation Output vs. Data



With *photometric-ID*,
Sim(Ia+CC) should
match observed HR

PS1 – Jones et al., 2018

Analysis Output → SNTABLES

- SNANA table before fit (1 row per SN)
- FITRES table after cuts+fit (1 row per SN)
- LCPLOT table with light curve & best-fit curve

Analysis Output → SNTABLES

- SNANA table before fit (1 row per SN)
- FITRES table after fit (1 row per SN)
- LC PLOT table with light curve & best-fit curve

Formats:

- TEXT (1 file per table)
- HBOOK (all tables → 1 file)
- ROOT (all tables → 1 file)

- Easier to install SNANA without HBOOK or ROOT
(but will miss most output variables from analysis)
- New format can be added if interface routines are provided

Analysis Output → SNTABLES

- SNANA table before fit (1 row per SN)
- FITRES table after fit (1 row per SN)

- ALL variables stored in **HBOOK & ROOT** format
(compared to TEXT, more efficient to read & write)
- Subset stored in **TEXT** for input to cosmology fit
- Utility to append **TEXT** file (from HBOOK or ROOT)
- See “sntable_dump” to view, dump, append tables
- Slowly moving away **TEXT** (except for debug)

Putting it all Together in a Photometric-SNIa-Cosmology Analysis

SNANA

- Simulate SNIa biasCor
- Simulate SNCC training
- Light curve fitting (data+sim)
- Beams with Bias Corr (BBC)
- Create COV(stat+syst)

Outside SNANA

- Classification
- Cosmology fit

Putting it all Together in a Photometric-SNIa-Cosmology Analysis

SNANA

- Simulate SNIa biasCor
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Outside SNANA

- Classification
- Cosmology fit

“Pippin” glues it all together:

<https://github.com/Samreay/Pippin>

Putting it all Together in a Photometric-SNIa-Cosmology Analysis

SNN: Moller and Boissiere, arXiv:1901.06384
SNIRF: Dai et al, arXiv:1701.05689
(more classifiers can be added)

Outside SNANA

- Classification
- Cosmology fit

CosmoMC: Lewis and Bridle, arXiv:astro-ph/0205436

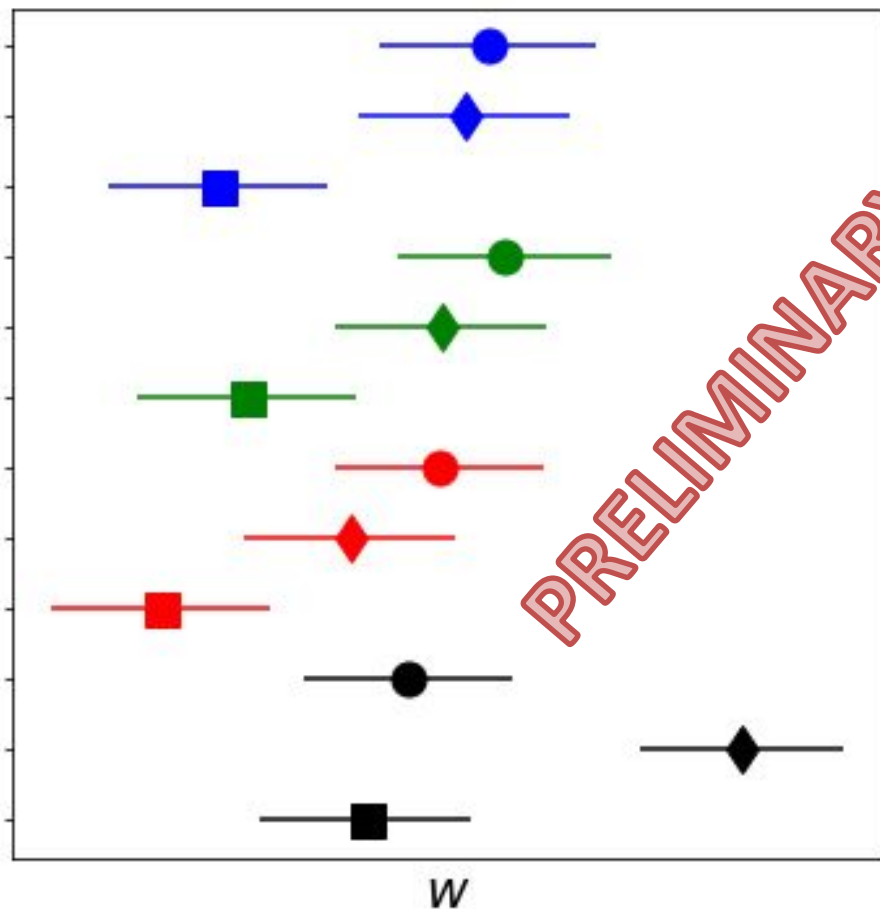
“Pippin” glues it all together:

<https://github.com/Samreay/Pippin>

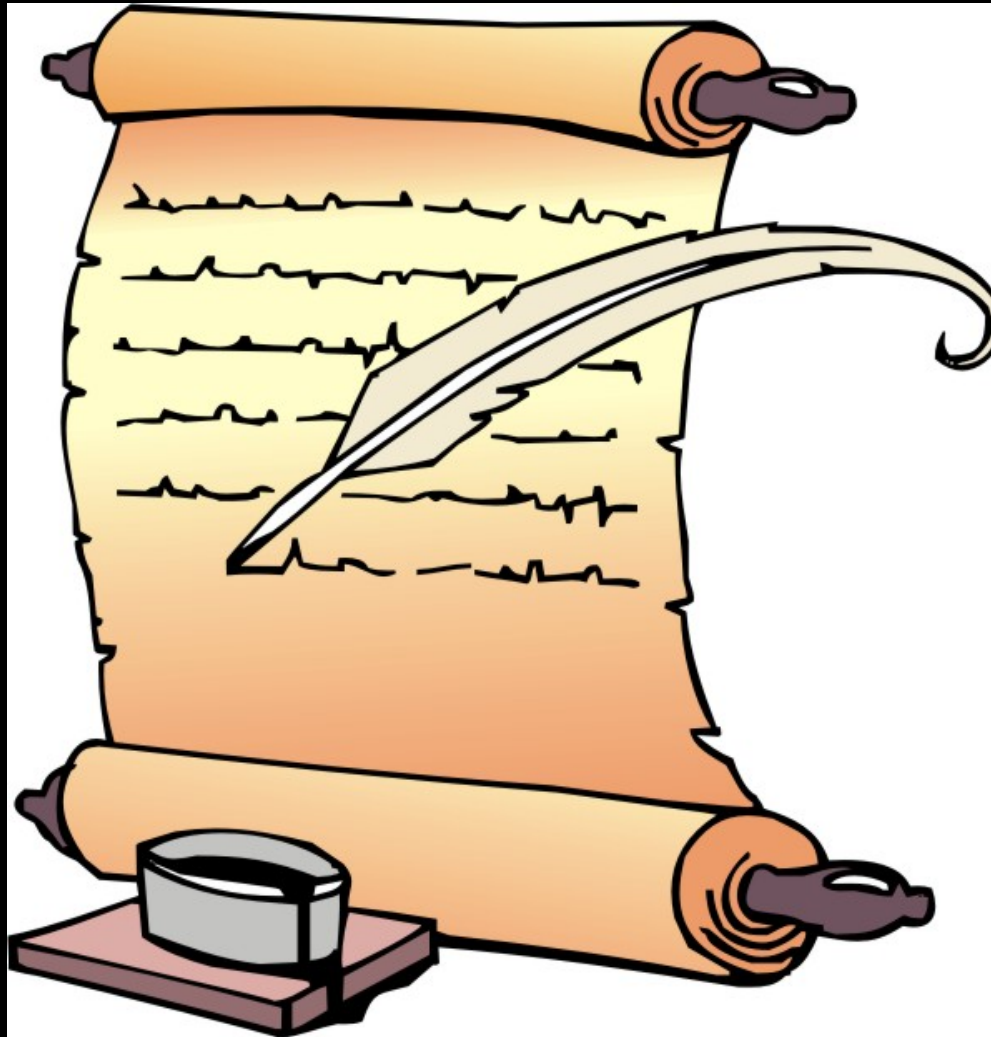
Power of Pippin: Illustration with Simulated DES+LOWZ

Classifier x simCC(training) x simCC(BBC prior)

SNN trained on V19; CCprior map:V19
SNN trained on V19; CCprior map:J17
SNN trained on V19; CCprior map:H11
SNN trained on J17; CCprior map:V19
SNN trained on J17; CCprior map:J17
SNN trained on J17; CCprior map:H11
SNN trained on DESCC; CCprior map:V19
SNN trained on DESCC; CCprior map:J17
SNN trained on DESCC; CCprior map:H11
Perfect Classifier; CCprior map:V19
Perfect Classifier; CCprior map:J17
Perfect Classifier; CCprior map:H11



SNANA Documentation



SNANA Documentation

- Learning SNANA is like learning to fly a plane; forget the manual, you need experienced person in the cockpit with you.

Why did my job abort ?

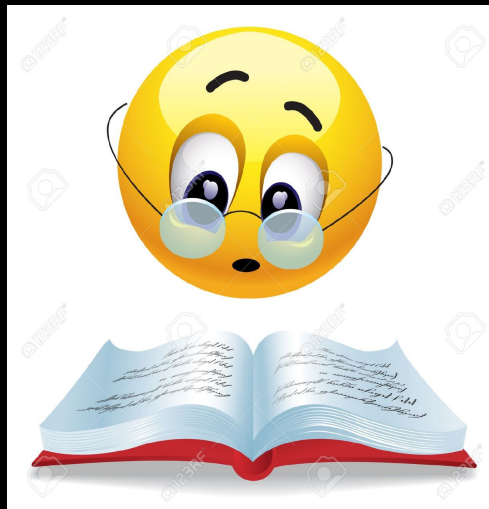
```
LC: 2000
ION: CODETEST_SALT2_SDSS_TEXT_XXX
ICE: RANDOM
ICE: SALT2_Guy10
E_INCLUDE: $SNANA_TESTS/inputs/SIMGEN_INCLUDE_SALT2.Input
FILE: $SNANA_ROOT/simlib/SDSS/SDSS2005_ugriz_SIMLIB
Ia: 120 106
KCOR_FILE: $SNANA_ROOT/kcor/SDSS/SDSS_D012010/kcor_SDSS_Bessel190_B017.fits
GENMAG_SMEAR_MODELNAME: G10
SEARCHEFF_PIPELINE_LOGIC_FILE: $SNANA_ROOT/models/searcheff/SEARCHEFF_PIPELINE_LOGIC.DAT
SEARCHEFF_PIPELINE_EFF_FILE: $SNANA_ROOT/models/searcheff/SEARCHEFF_PIPELINE_SDSS.DAT
SEARCHEFF_SPEC_FILE: $SNANA_ROOT/models/searcheff/SEARCHEFF_SPEC_SDSS.DAT
SEARCHEFF_HOST_FILE: $SNANA_ROOT/models/searcheff/SEARCHEFF_HOST_SDSS.DAT
APPLY_SEARCHEFF_OPT: 3
# selection criteria for generation
GENFILTERS: ugriz
GENRANGE_RA: -50.0 +59.0 deg
GENRANGE_DECL: -1.258 +1.258 deg
GENRANGE_PEAKRJD: 53625.0 53685.0
GENSIGMA_SEARCH_PEAKRJD: 1.0
GENRANGE_REDSHIFT: 0.02 0.45
GENSIGMA_REDSHIFT: 0.0012
GENRANGE_TREST: -20.0 80.0 # rest epoch relative to peak (days)
RANSEED: 128473 # random number seed
# smear flags: 0=off, 1=on
SMEARFLAG_FLUX: 1 # photo-stat smearing of signal, sky, etc ...
SMEARFLAG_ZEROPT: 1 # smear zero-point with zptsig
DNDZ: POWERLAW 2.6E-5 1.5 # SN rate ~ (1+z)^1.5
FORMAT_MASK: 2 1 2=terse/text 32=FITS 1=verbose
SIMGEN_DUMP: 14 CID GENTYPE SNTYPE GENZ DLMAG S2x1 S2c
SNRMAX SNRMAX2 SNRMAX3
SNRMAX_g SNRMAX_r SNRMAX_i SNRMAX_z
```

What does the red engine light mean ?

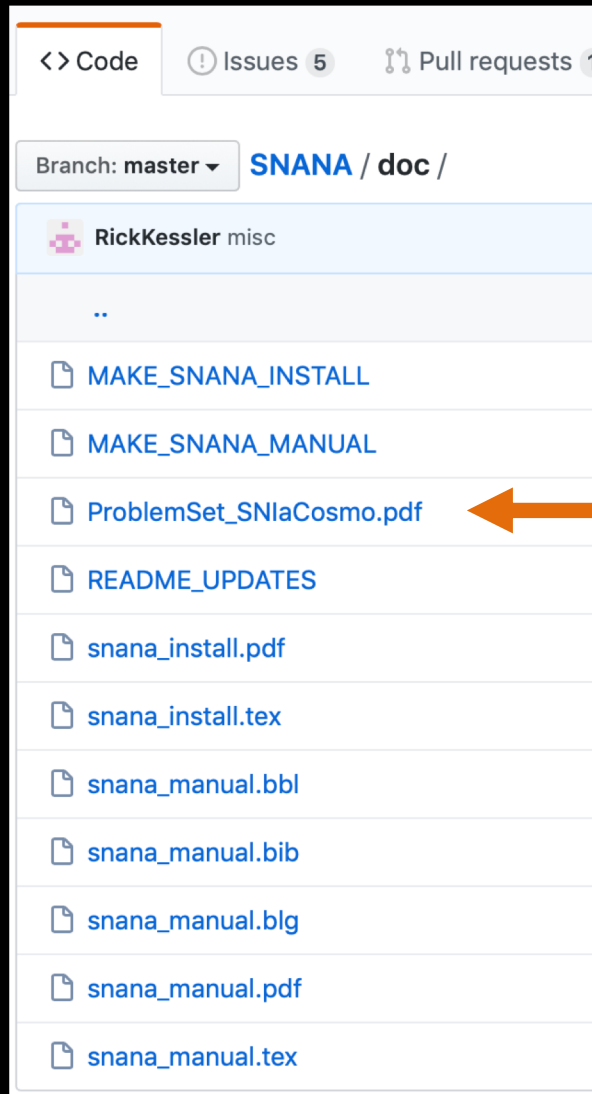


Manual

- 180 page manual with no interface except 'preview'
- Reading manual not so helpful, but recommend skimming table-of-contents



For Beginners and Profs looking for Homework Problems



Tracking Changes

```
des20.fnal.gov> tail -50 $SNANA_DIR/doc/README_UPDATES
```

```
=====
```

```
v10_42f (Feb 8 2016)
```

```
***** IMPORTANT(v10_42f) *****
```

```
***** USEFUL(v10_42f) *****
```

```
***** MISCELLANEOUS(v10_42f) *****
```

Ignore boring “MISC”
that is mainly for me

```
sntools_output_root.c:
```

```
in SNTABLE_READPREP_ROOT, add missing return(NVAR) at end.  
This bug was tripped up by the recent -O1 optimization, but  
amazingly seemed to work on other machines
```

```
Add MJD to SNLCPAK tree, so that we don't have to use the  
clumsy method pf MJD = PEAKMJD+TOBS.
```

Tracking Changes

v10_42g (Feb 19 2016)

***** IMPORTANT(v10_42g) *****

SALT2mu.c : major refactor and update to implement BEAMS-like fit using simulation to define the CC prior.
See new inputs: simfile_ccprior and varname_p1a

***** USEFUL(v10_42g) *****

New plotting function \$SNANA_DIR/util/ovdatamc.py (by D.Jones) operates on ascii FITRES files from data and sim. Overlays simulation separately for SNIa and SNCC.

snana.car: new SNLCINP namelist SIMLIB_OUT = 'bla.simlib' will create simlib file from data.

***** MISCELLANEOUS(v10_42g) *****

snlc_sim.c, sntools_host.c:
New sim-input key
HOSTLIB_GALID_PRIORITY: 0 500000
to give priority to GALID range.

Fix minor bug in gen_AV() [found by D.Jones]

split_and_fit.pl : new key SALT2mu_SIMVERSION to pass simFile to SALT2mu program (for bias cor and CCprior).

You should read
**IMPORTANT &
USEFUL** updates

Miscellaneous

- Report bugs ASAP as Github issue; don't just hack a private fix for yourself.
- Report compilation warnings.
- Think about how you can contribute to SNANA.
- In papers, SNANA citation is not enough; ***also cite source of models, template data, galaxy catalogs, etc.***



Conclusion

From the SNANA Legal Team:

- **SIDE EFFECTS** include, but are not limited to: confusion, frustration, watery eyes, headaches, weight loss, weight gain, systematics-limited results, incorrect results, denial of tenure.
- Do not drive or operate heavy machinery while using SNANA.